## DETECTING THE MISSING BARYONS IN X-RAY ABSORPTION

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06/04 MARYLAND

#### MISSING BARYONS, IN THE FORM OF THE WHIM:

<u>Absorption</u>	<u>Imaging</u>	
Depend on background sources	No dependency on background sources	
Sensitive to density	Sensitive to density square	
No dependency on the XRB	Removal of the XRB	

<u>Ultraviolet</u>	<u>X-ray</u>
H I, O VI, C IV	O VII, O VIII, C V, Ne X
Plasma with T ~ 10 <sup>5</sup> -10 <sup>6</sup> K	Plasma with T > 10 <sup>6</sup> K

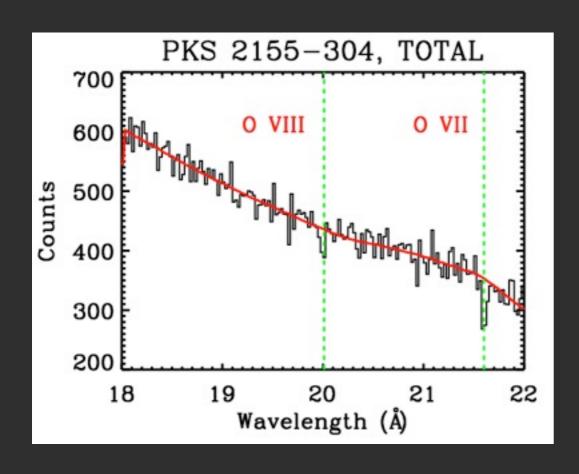
S5 0836+710 PKS 2149-306	Fang et al. (2001)	HETG	No detection
PKS 2155-304	Fang et al. (2002)	LETG	> 3-sigma
H 1821-643	Mathur et al. (2003)	LETG	2-sigma
3C 120	McKernan et al. (2003)	HETG	~ 3-sigma
LBQS 1228+1116	Fujimoto et al. (2004)	RGS	~ 2-3 sigma
Mkn 421	Nicastro et al. (2005)	LETG	> 3-sigma
X Comae	Takei et al. (2006)	RGS	2-sigma
PKS 0558-504	Nicastro et al. (2009)	RGS	~ 2-3 sigma

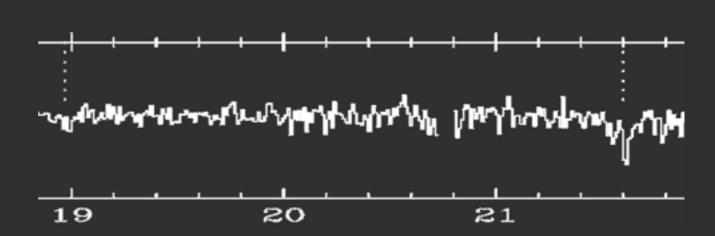
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#### ISSUE 1: INCONSISTENCY BETWEEN INSTRUMENTS

#### CHANDRA LETG

#### XMM RGS





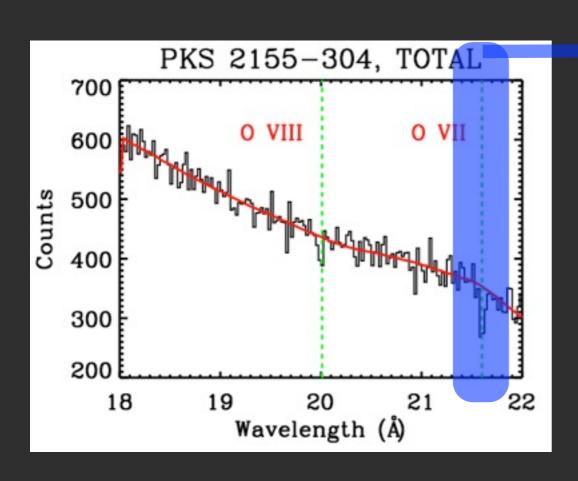
FANG ET AL. (2007)

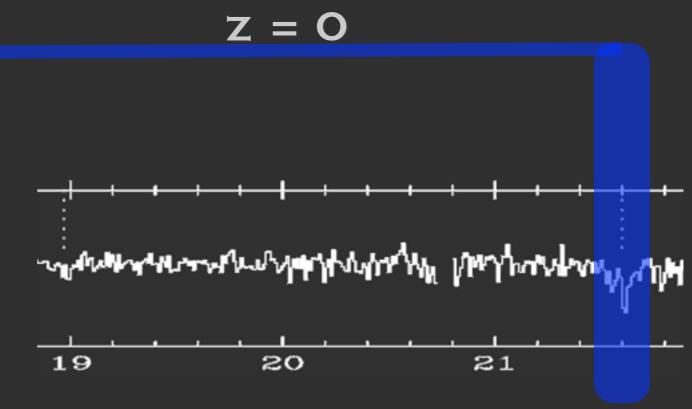
RASMUSSEN ET AL. (2003)

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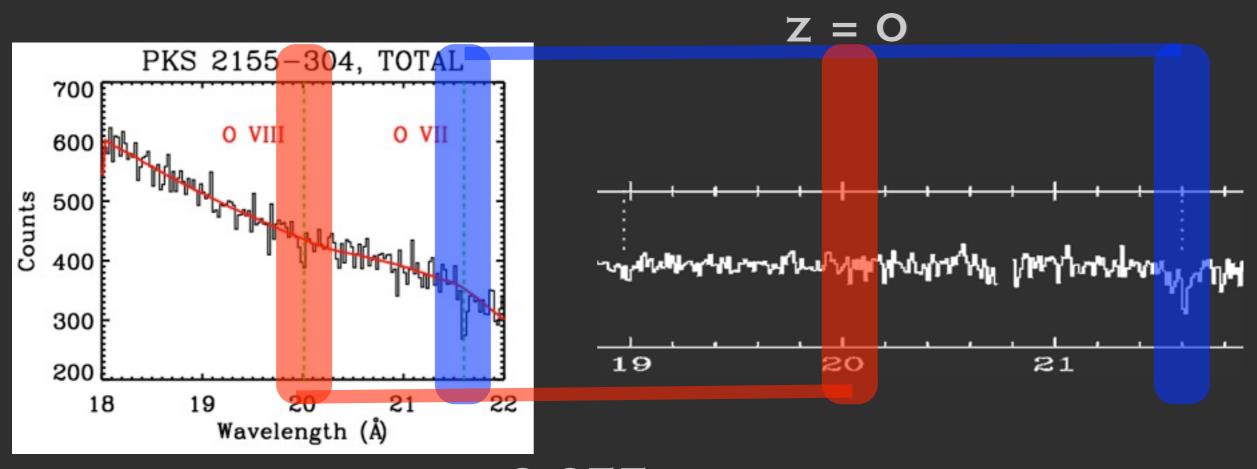
FANG ET AL. (2007)

RASMUSSEN ET AL. (2003)

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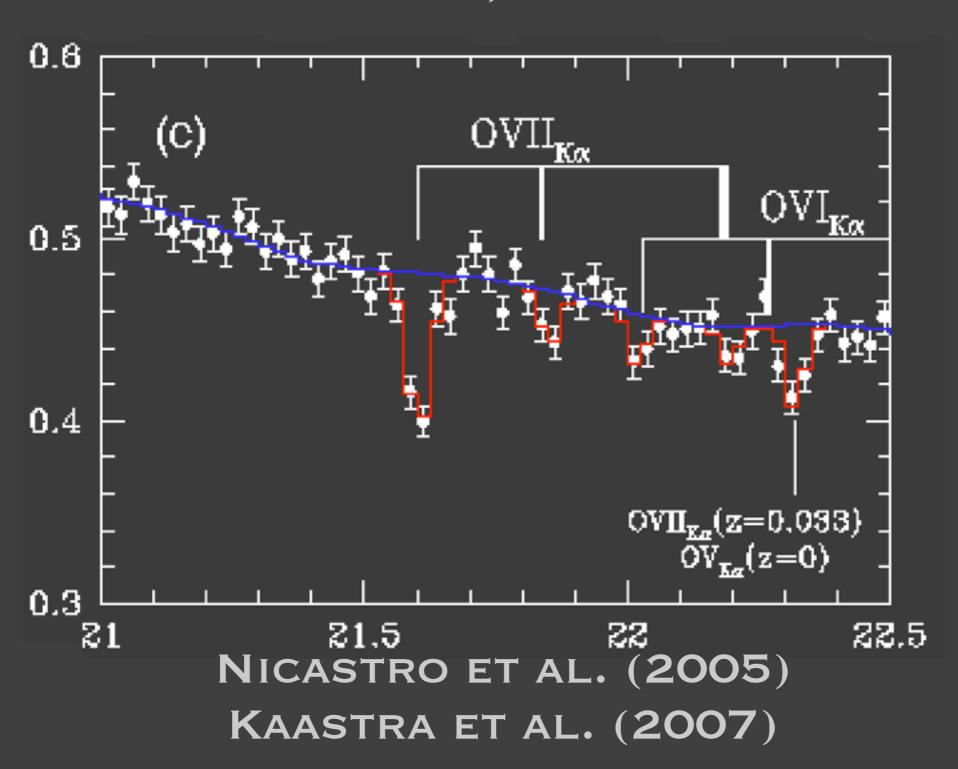
z = 0.055

FANG ET AL. (2007)

RASMUSSEN ET AL. (2003)

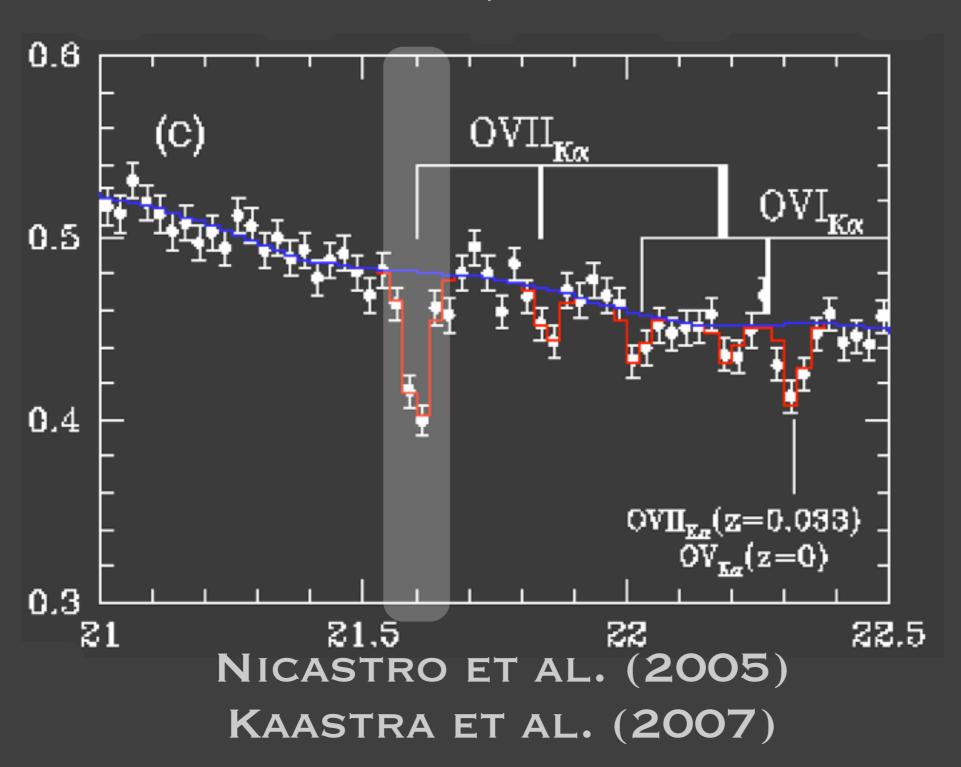
#### ISSUE 2: STATISTICS

MKN 421, CHANDRA



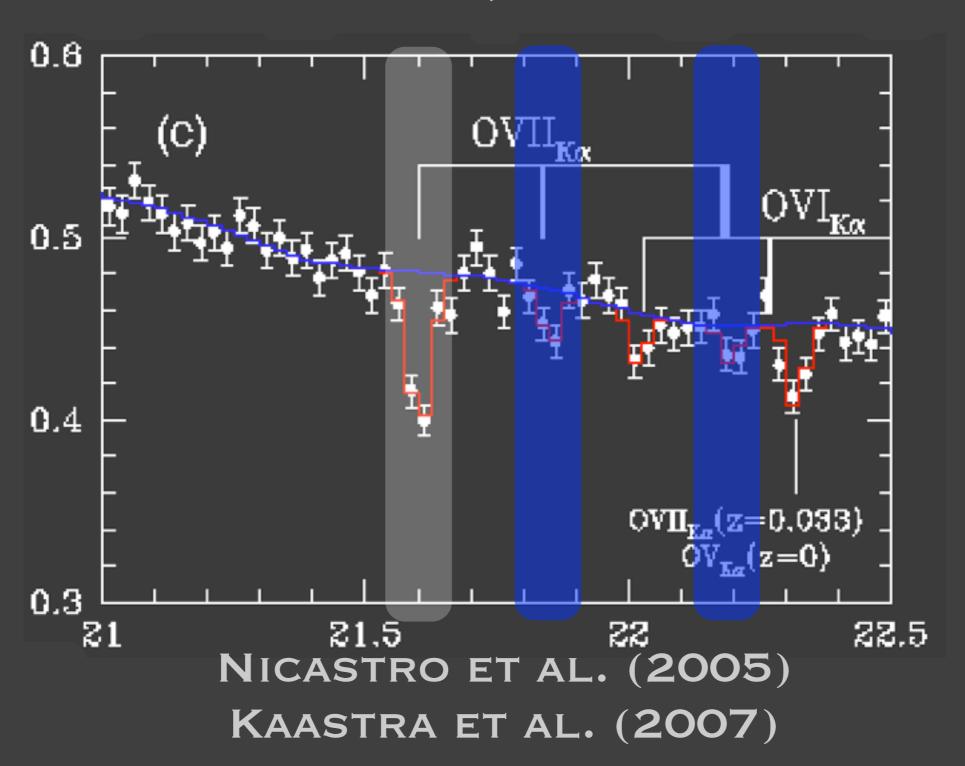
#### ISSUE 2: STATISTICS

MKN 421, CHANDRA



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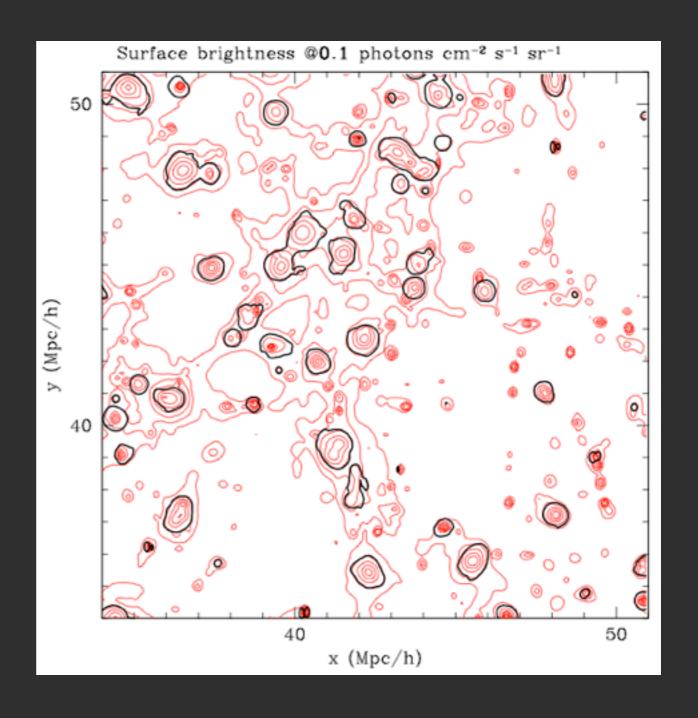
MKN 421, CHANDRA



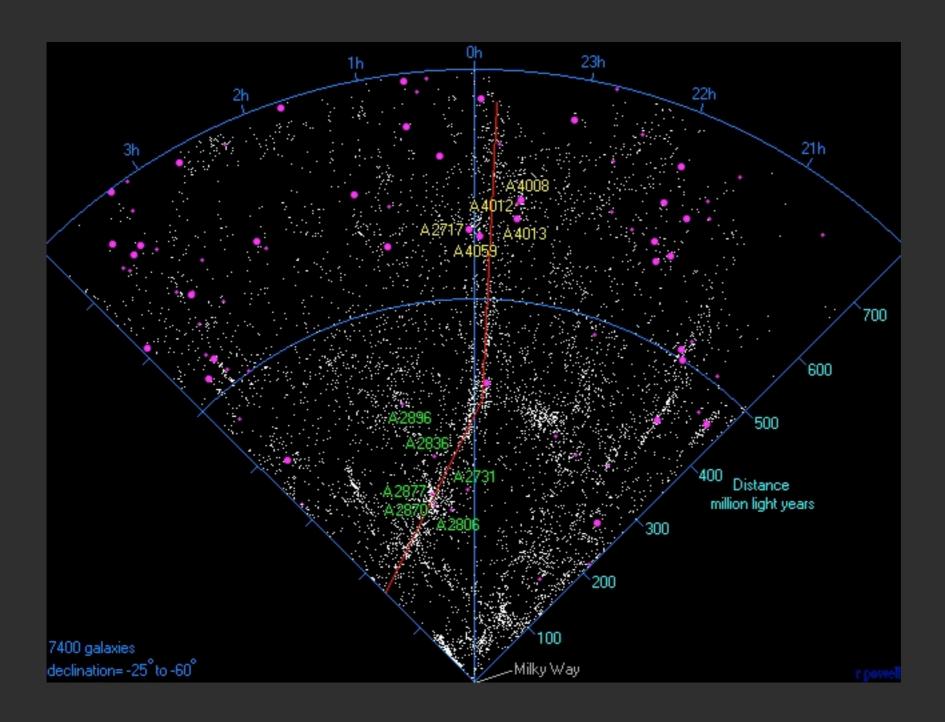
#### SOLUTION:

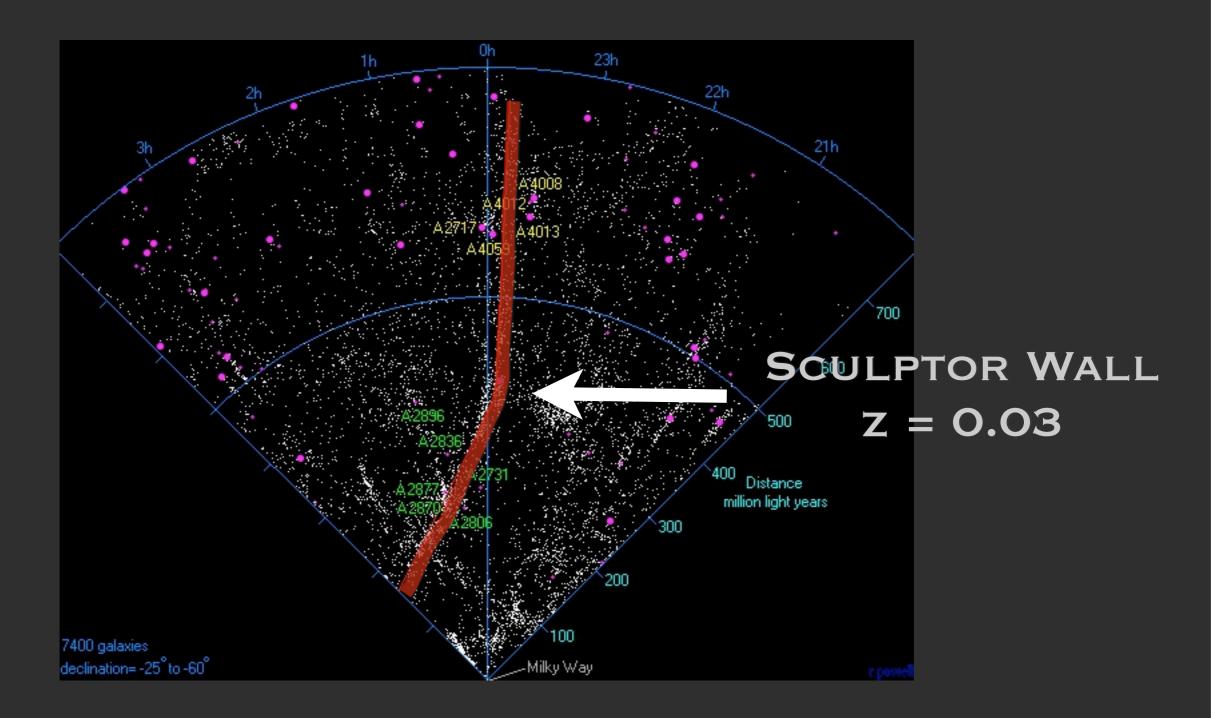
JOINT
DETECTION WITH
CHANDRA AND
XMM



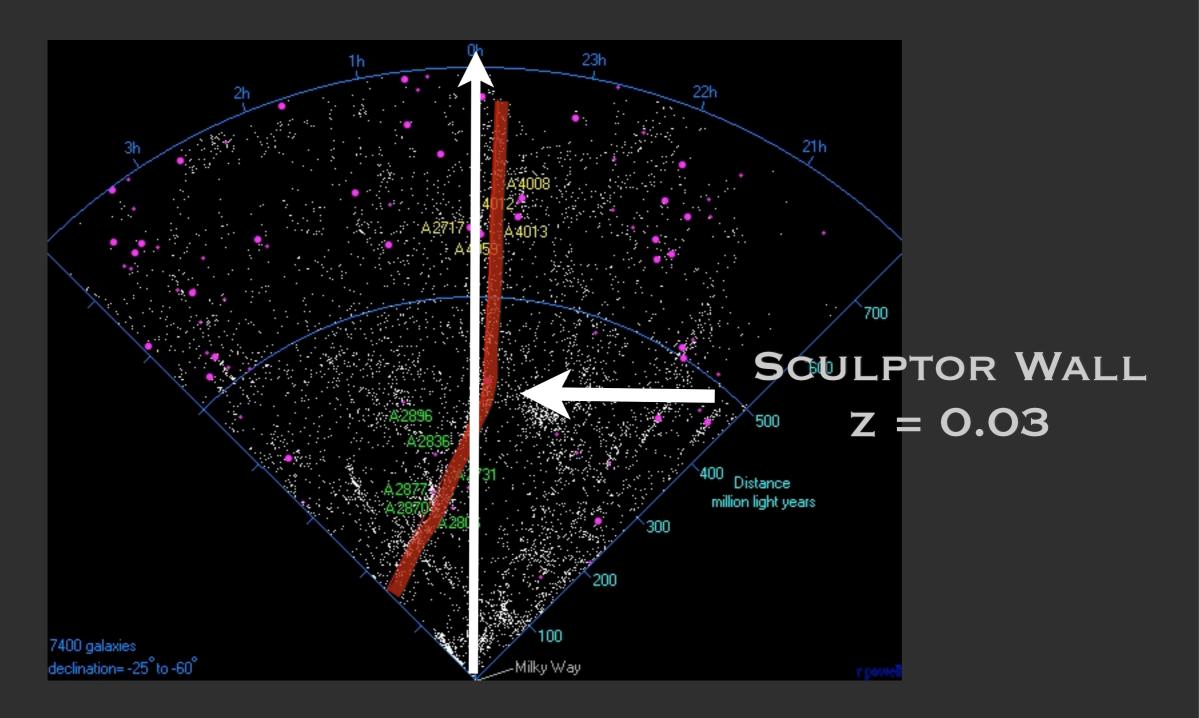


CEN & FANG (2006)

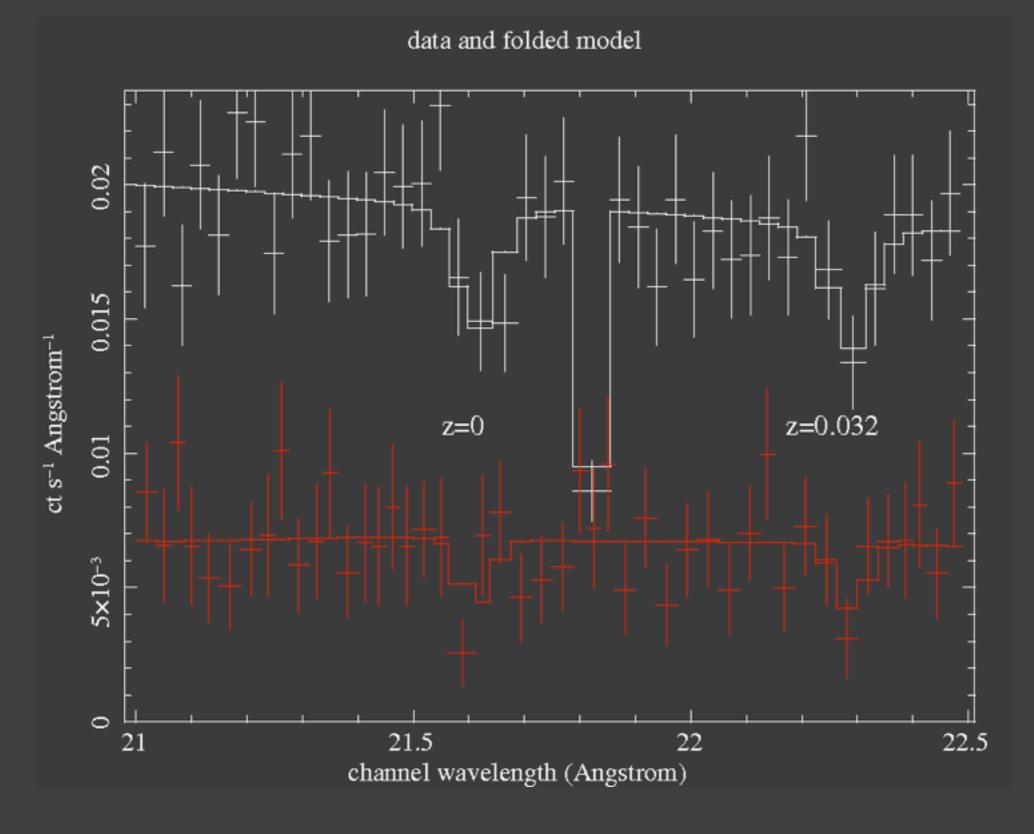




#### H 2356-309

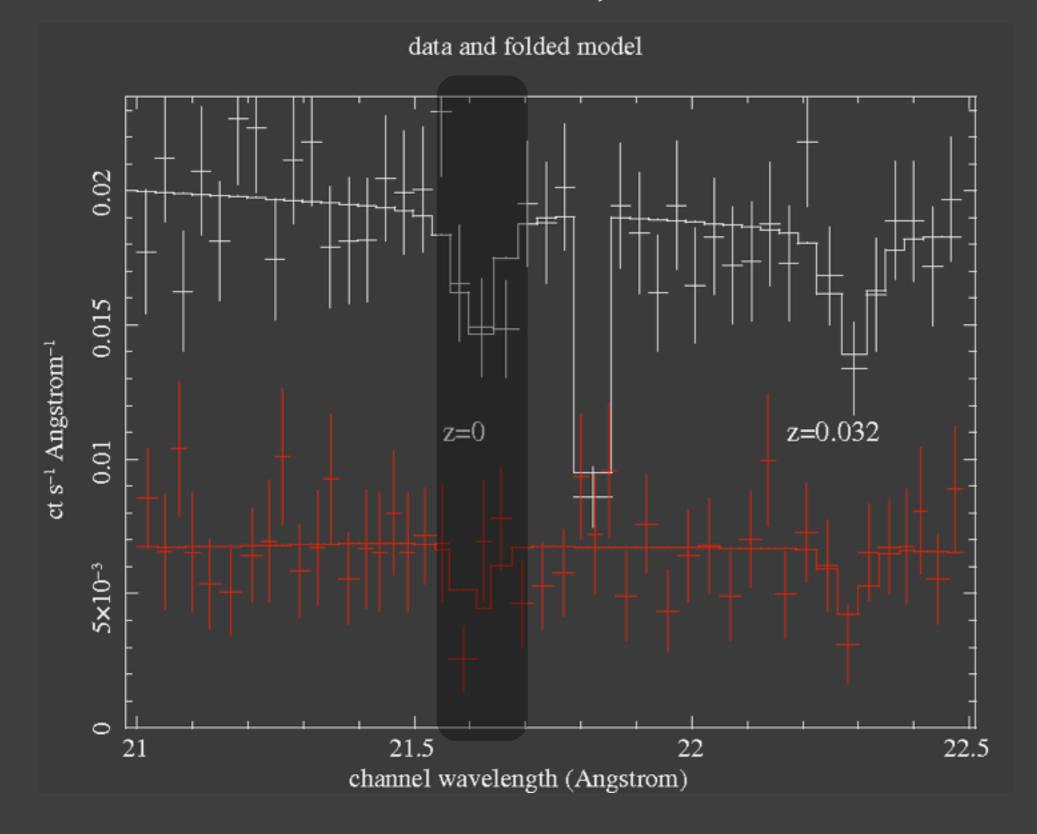


#### CHANDRA: 100 KSEC; XMM RGS: 130 KSEC



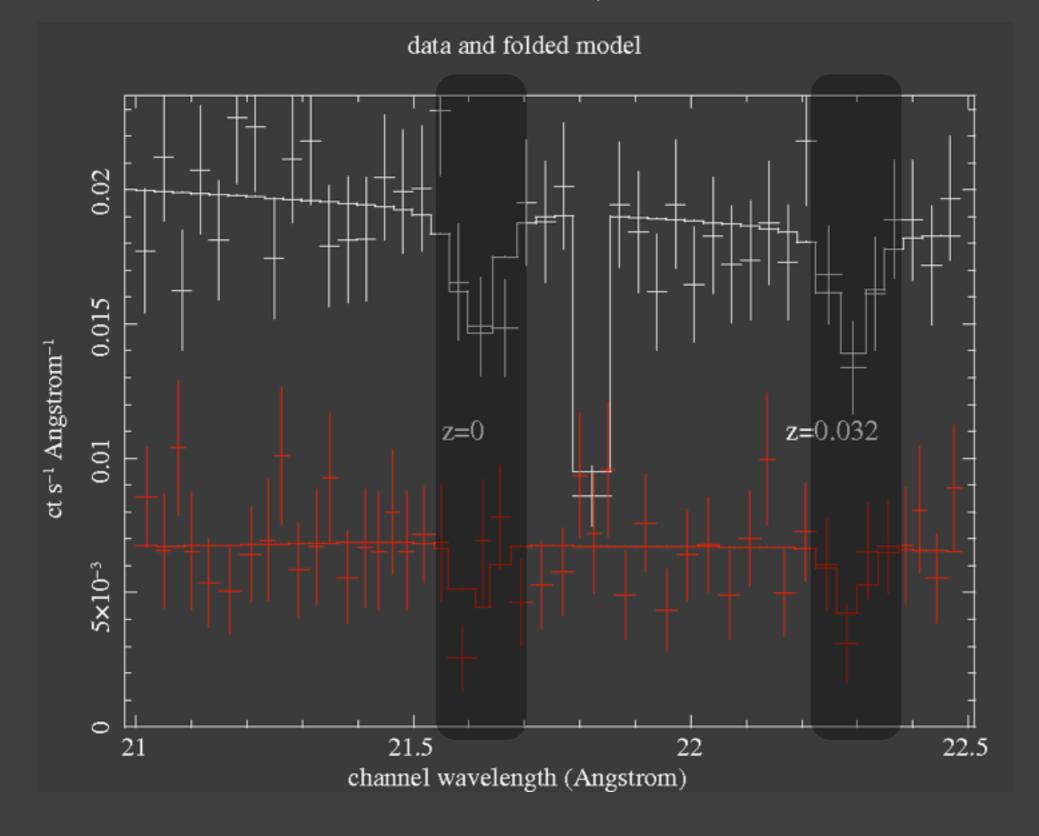
BUOTE ET AL. (2009)

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BUOTE ET AL. (2009)

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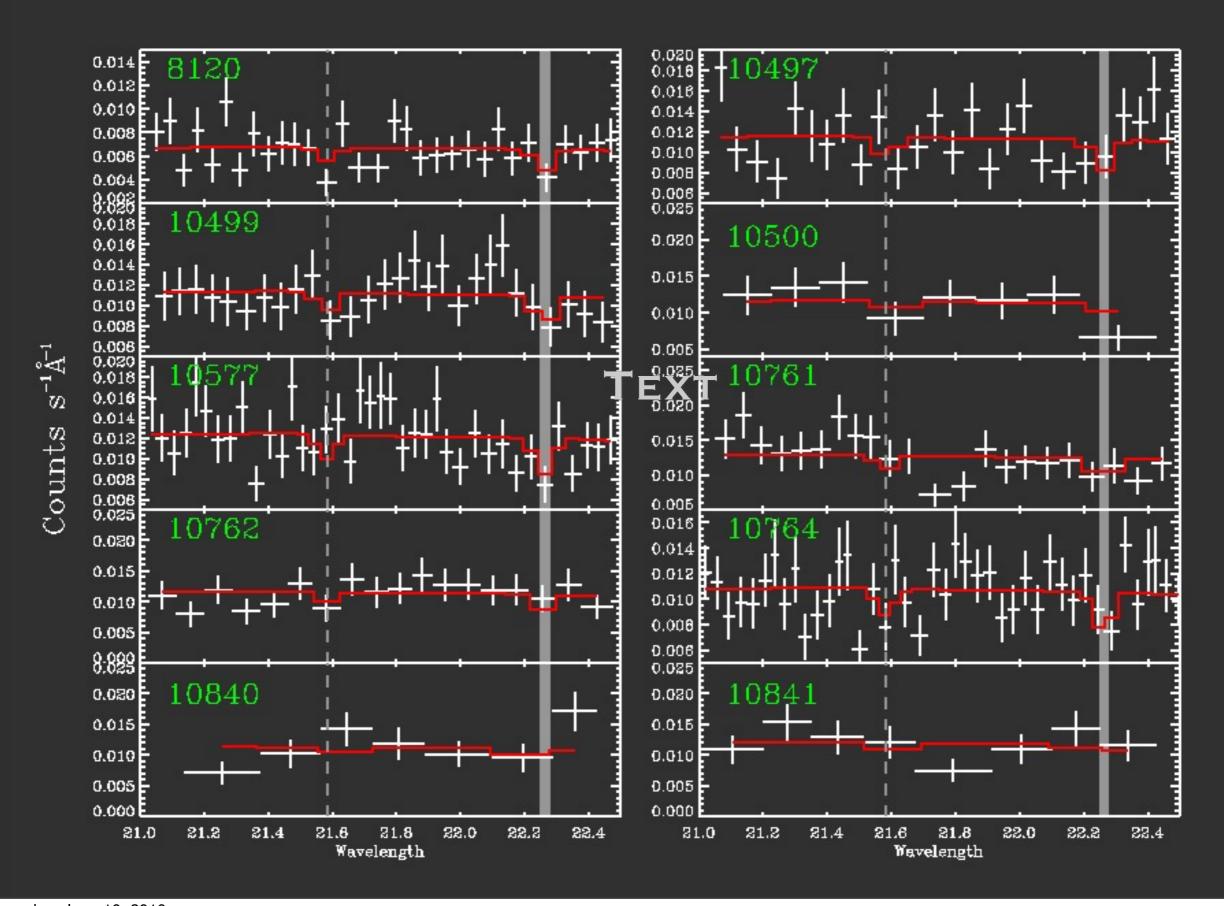


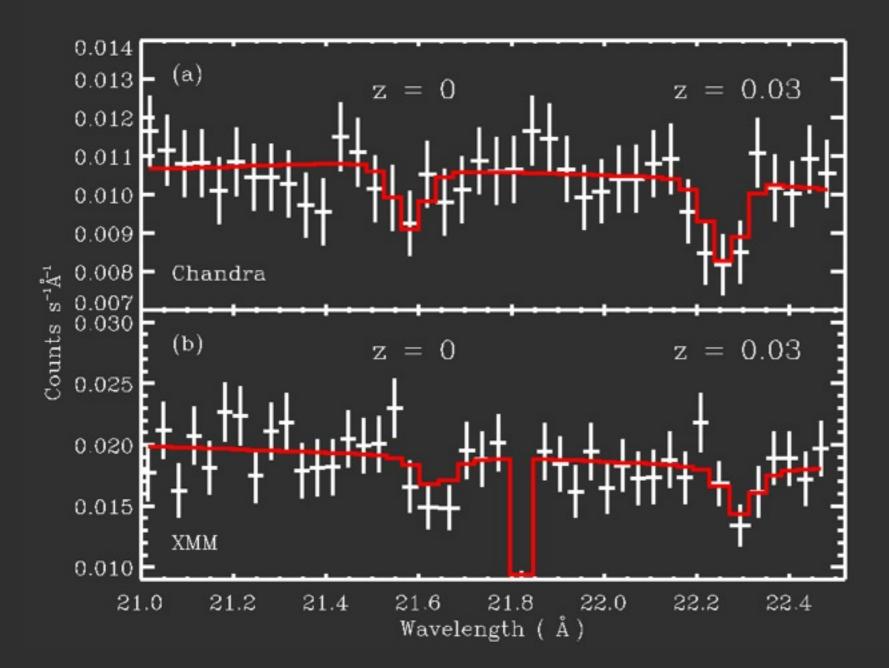
XMM 2.4-SIGMA

CHANDRA 1.7-SIGMA

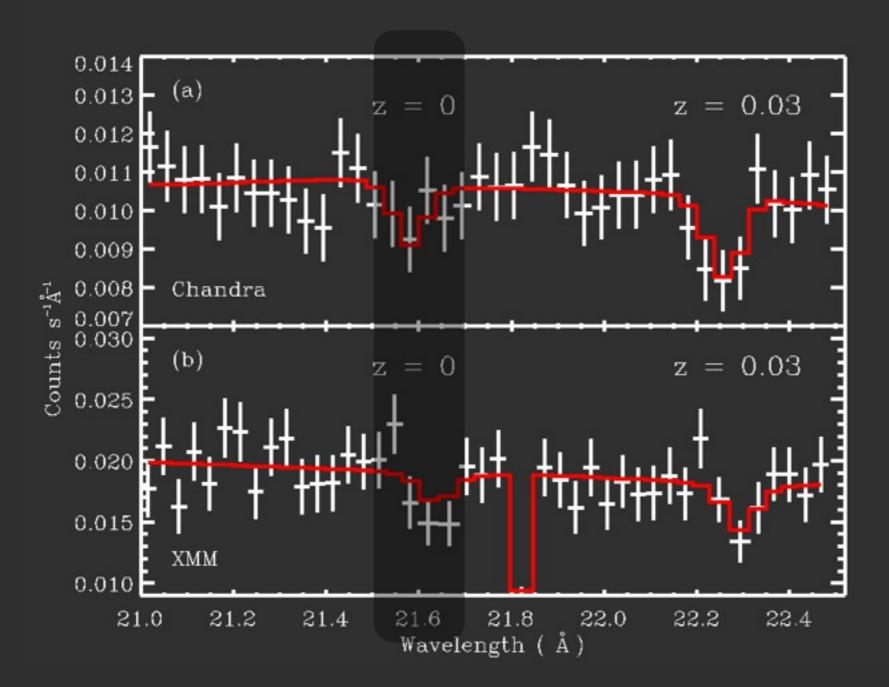
BUOTE ET AL. (2009)

#### CHANDRA 500 KSEC OBSERVATION

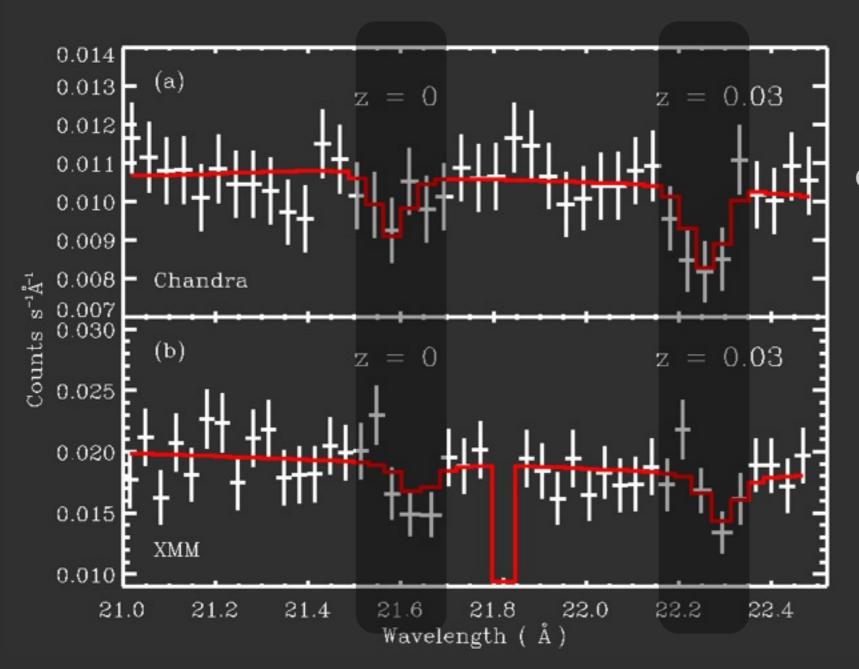




FANG ET AL. (2010)



FANG ET AL. (2010)

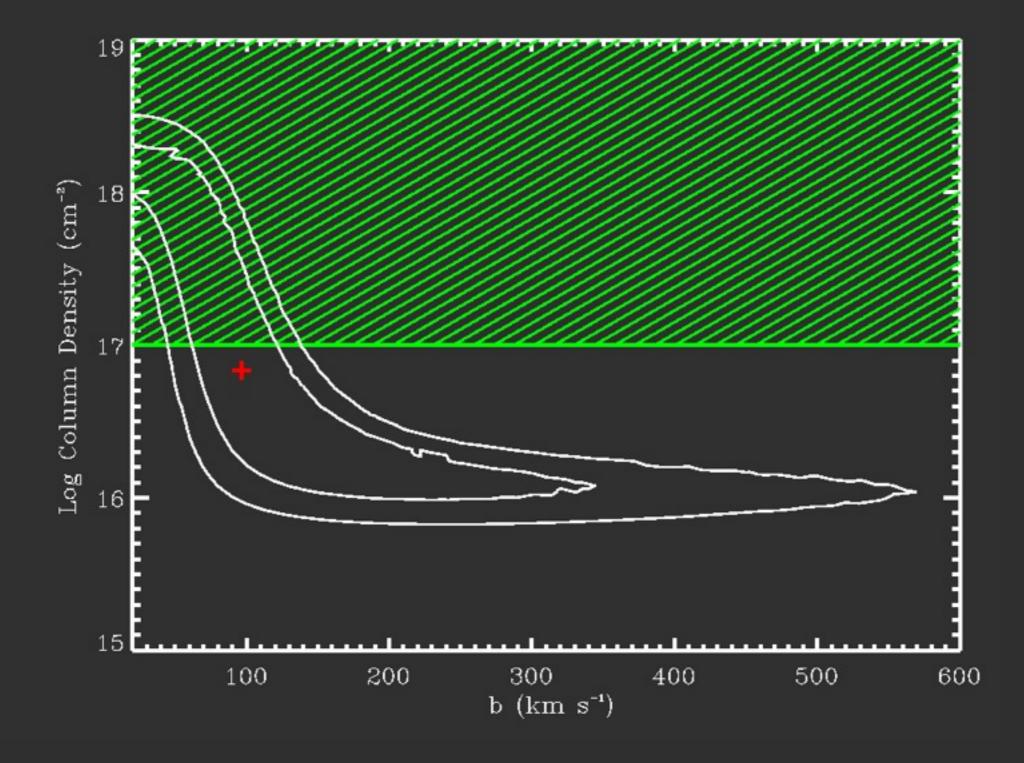


CHANDRA, 3.4-SIGMA

XMM, 2.4-SIGMA

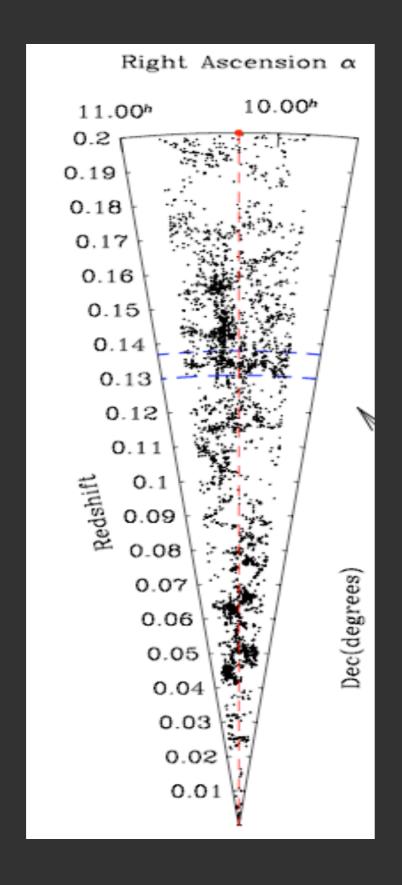
FANG ET AL. (2010)

# COLUMN DENSITY

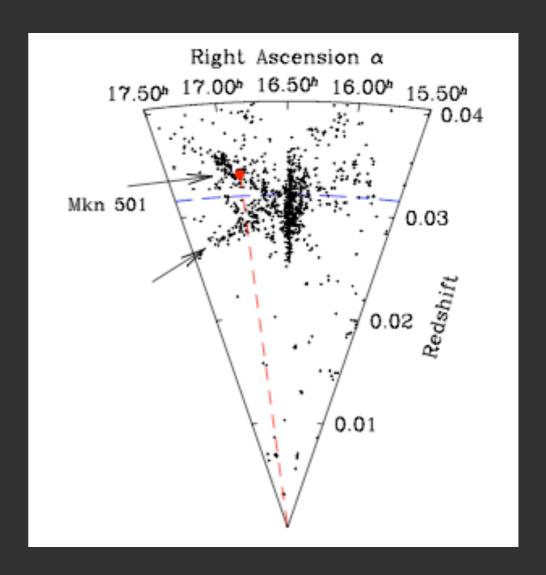


#### **DOPPLER-B PARAMETER**

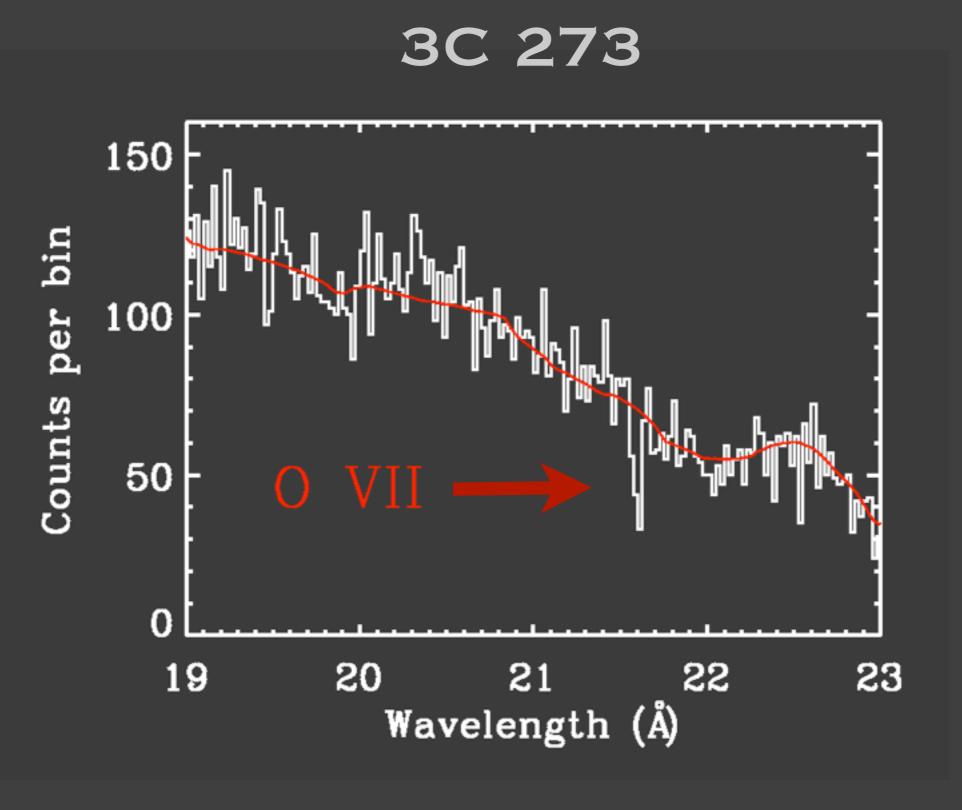
#### 1ES1011+496



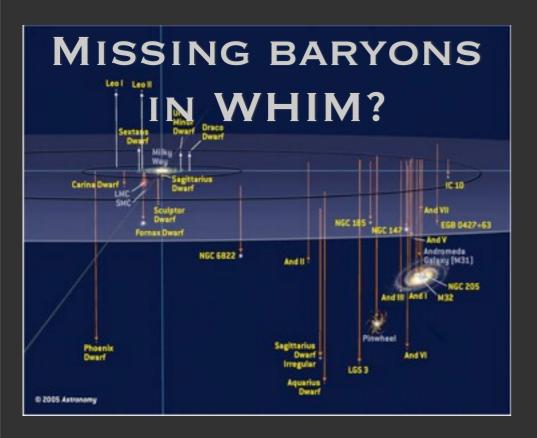
#### MKN 501



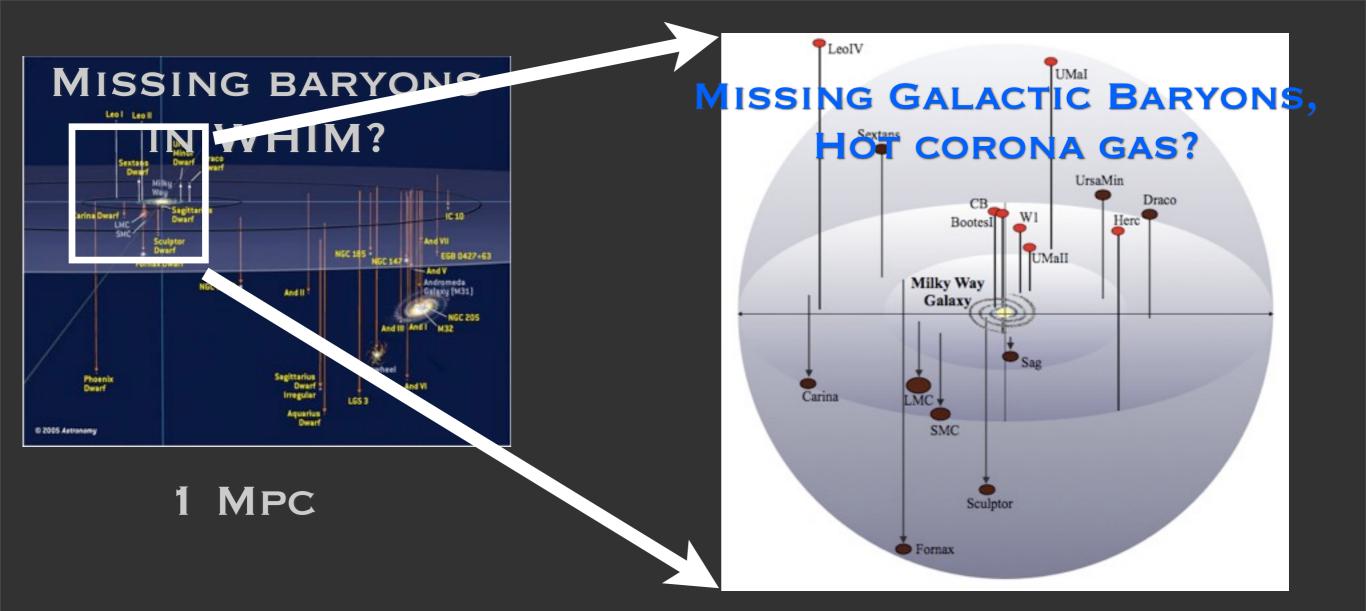
(APPROVED XMM 150 KSEC)



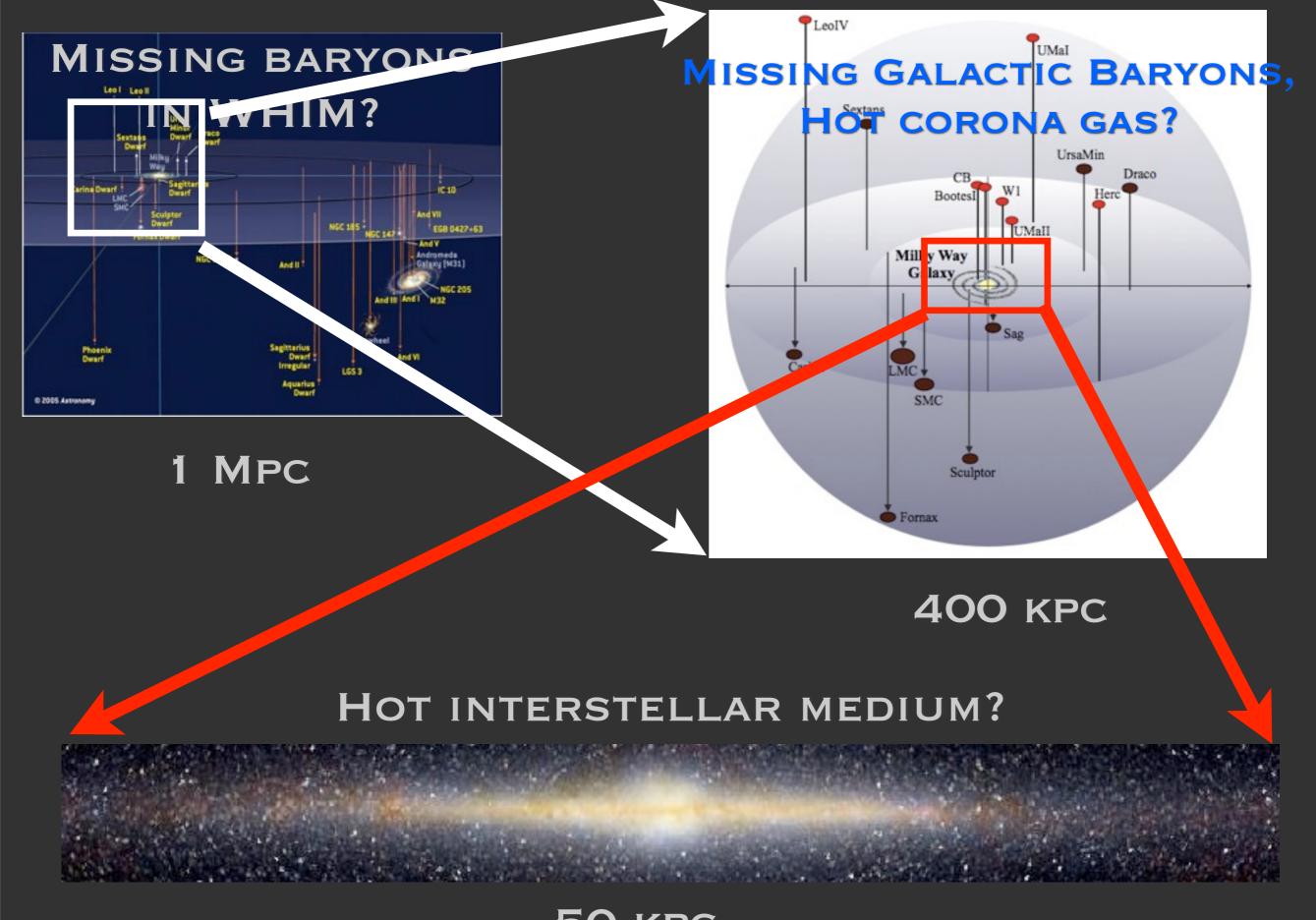
FANG ET AL. (2003)



1 MPC

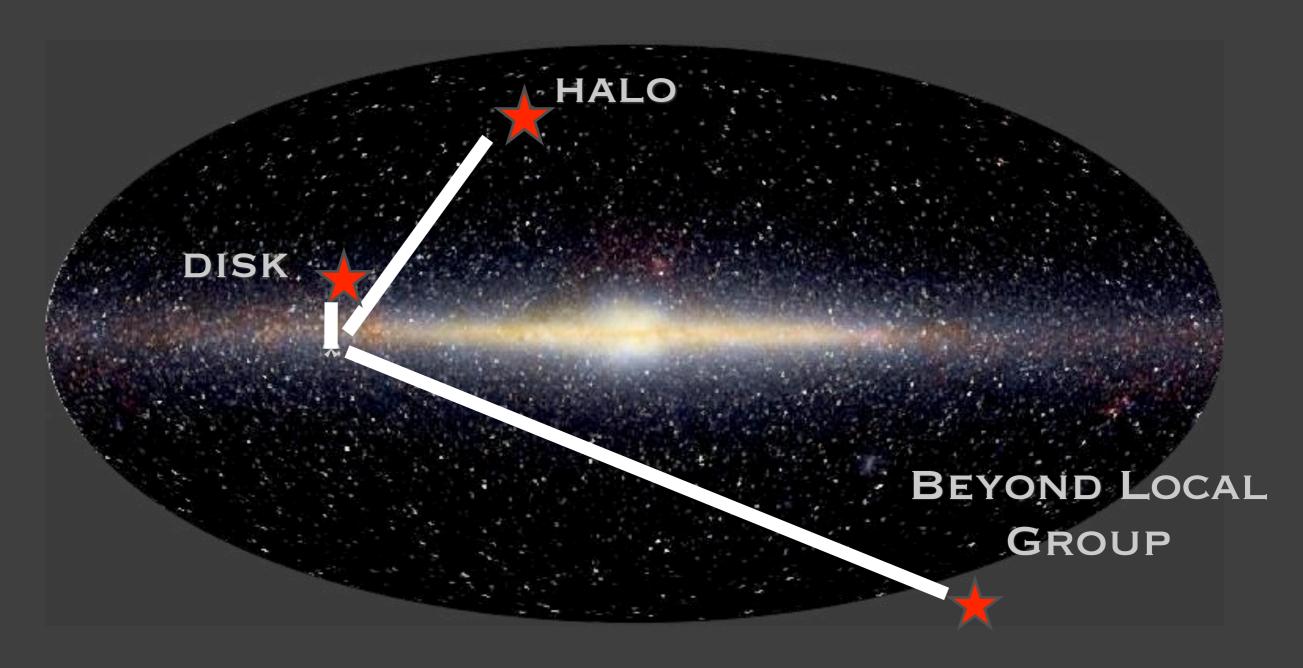


400 KPC



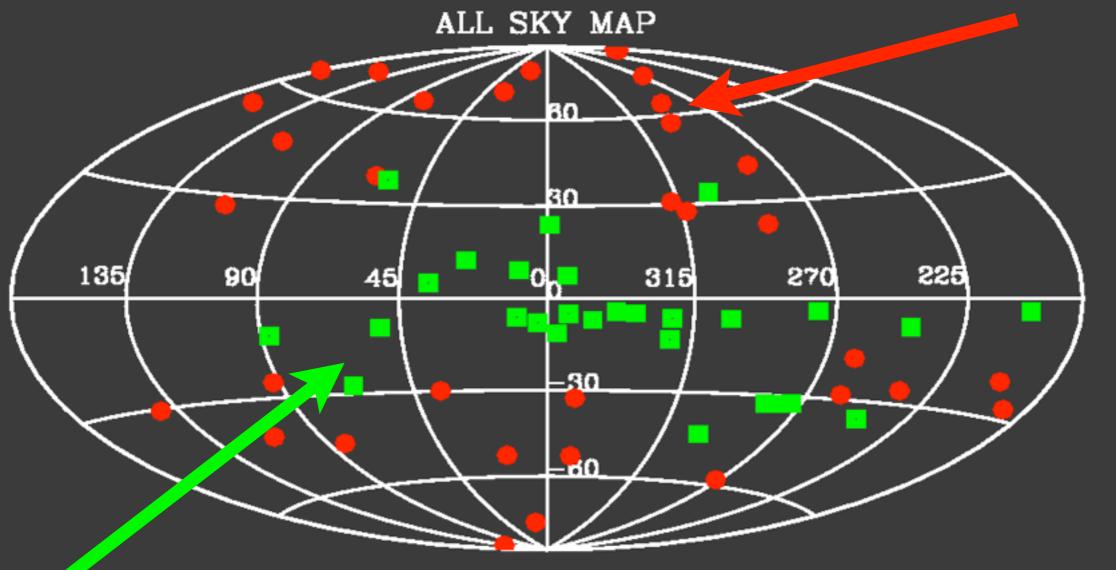
**50 KPC** 

### HOT GAS AROUND THE MILKY WAY: A DIFFERENTIAL DIAGNOSIS



FANG ET AL. (2010)



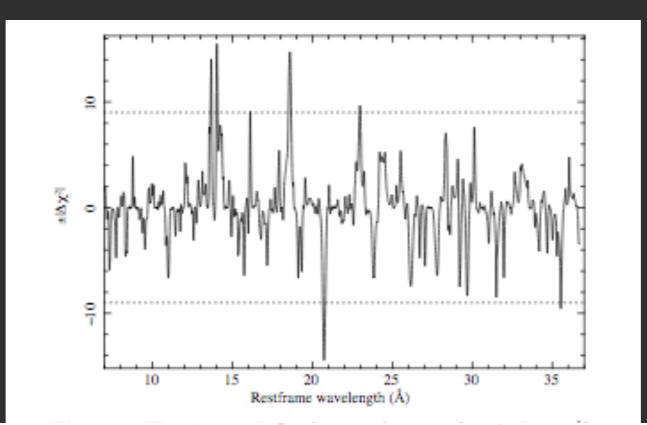


GALACTIC SOURCES

34 EXTRAGALACTIC (RED)

28 GALACTIC (GREEN), DISTANCE=[0.3, 50] KPC

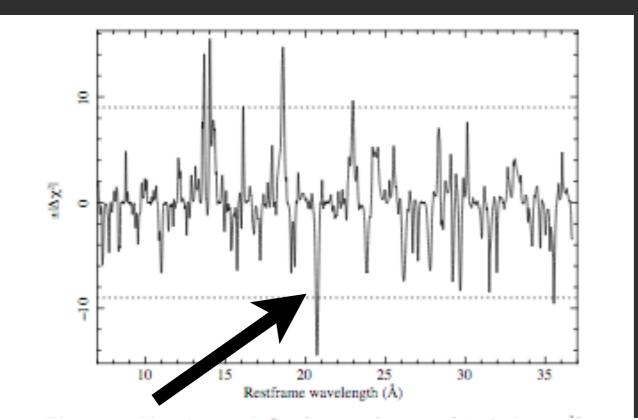
#### SOMETHING INTERESTING ...



cannot find a consistent interpretation for them all. There is no plausible identification for the possible absorption line at ~20.7 Å. In general, the soft X-ray spectrum contains no evidence of a warm absorber type outflow.

#### BY WHOSE-NAME-SHOULD-NOT-BE-MENTIONED, 2008

#### SOMETHING INTERESTING ...



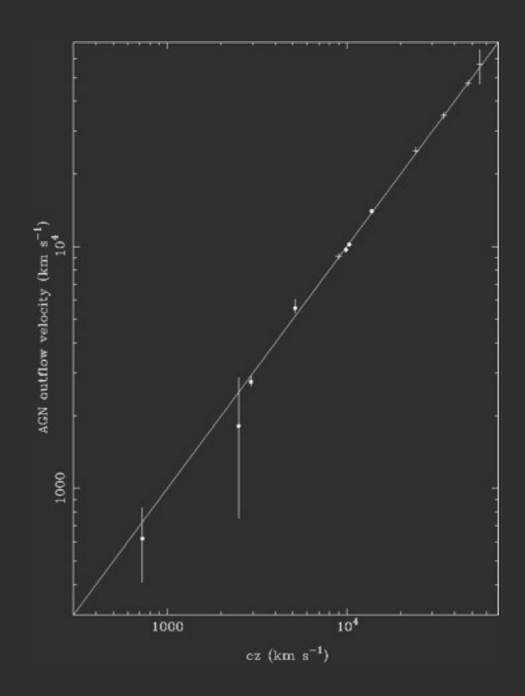
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#### BY WHOSE-NAME-SHOULD-NOT-BE-MENTIONED, 2008

#### AND SOMETHING LIKE A "COSMIC CONSPIRACY"?

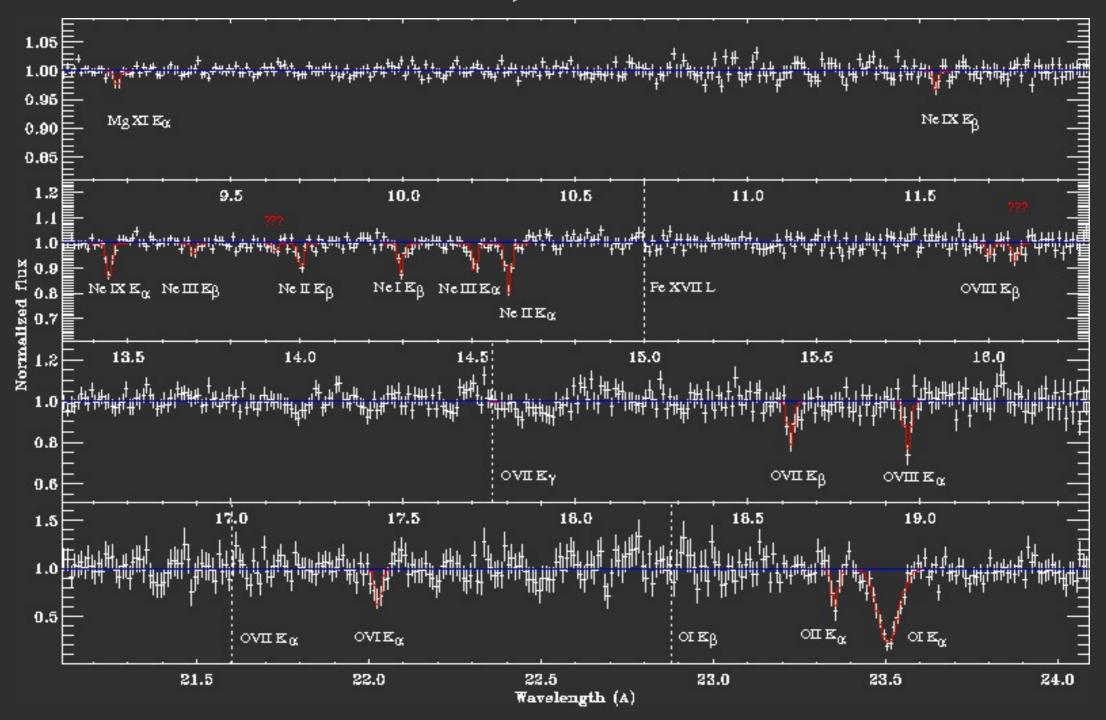
AGN
OUTFLOW
OR LOCAL
HOT GAS?



McKernan et al. (2003)

#### AND SOMETHING WEIRD ...

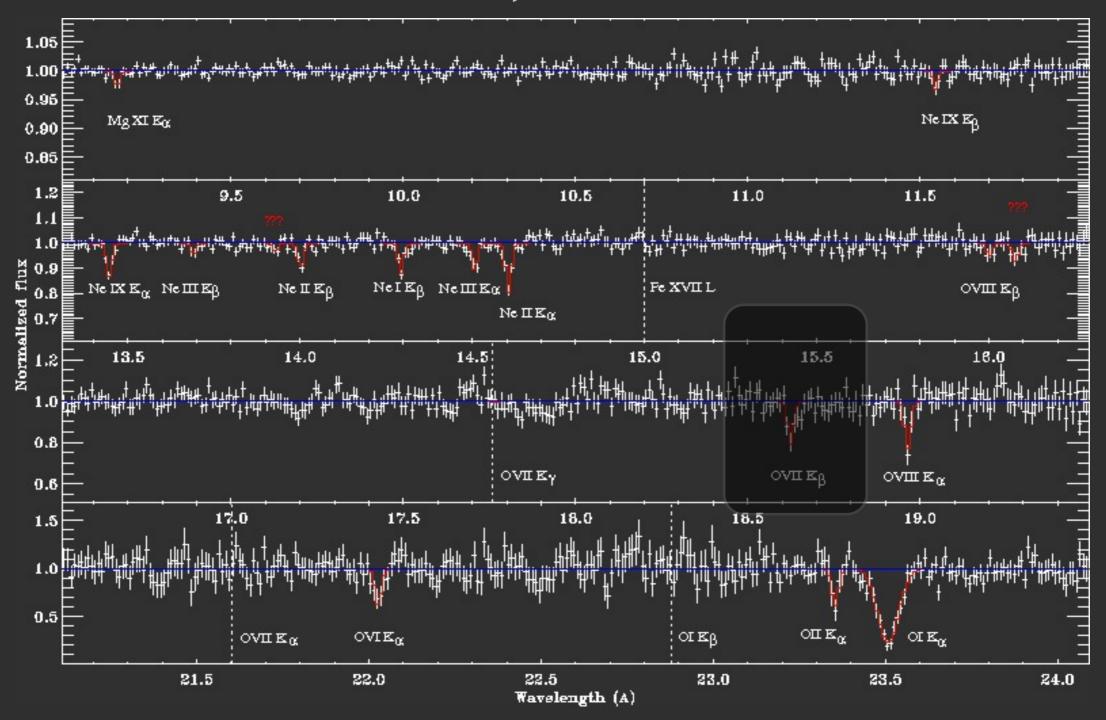
#### CYG X-2, CHANDRA



#### YAO ET AL. (2009)

#### AND SOMETHING WEIRD ...

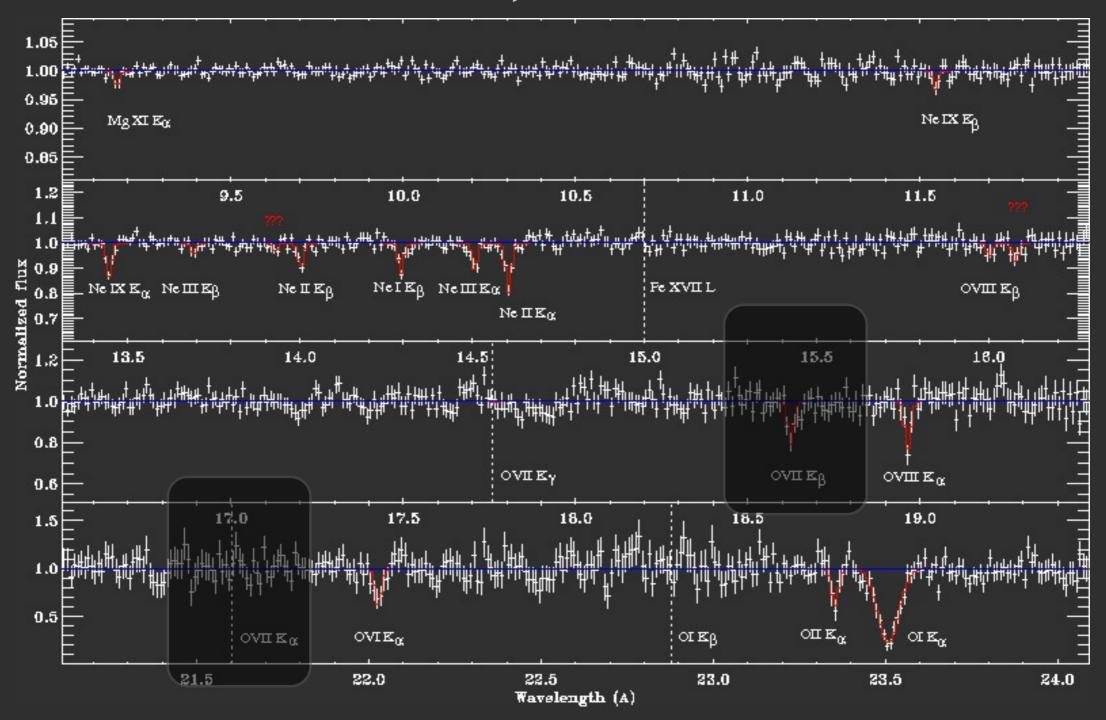
#### CYG X-2, CHANDRA



#### YAO ET AL. (2009)

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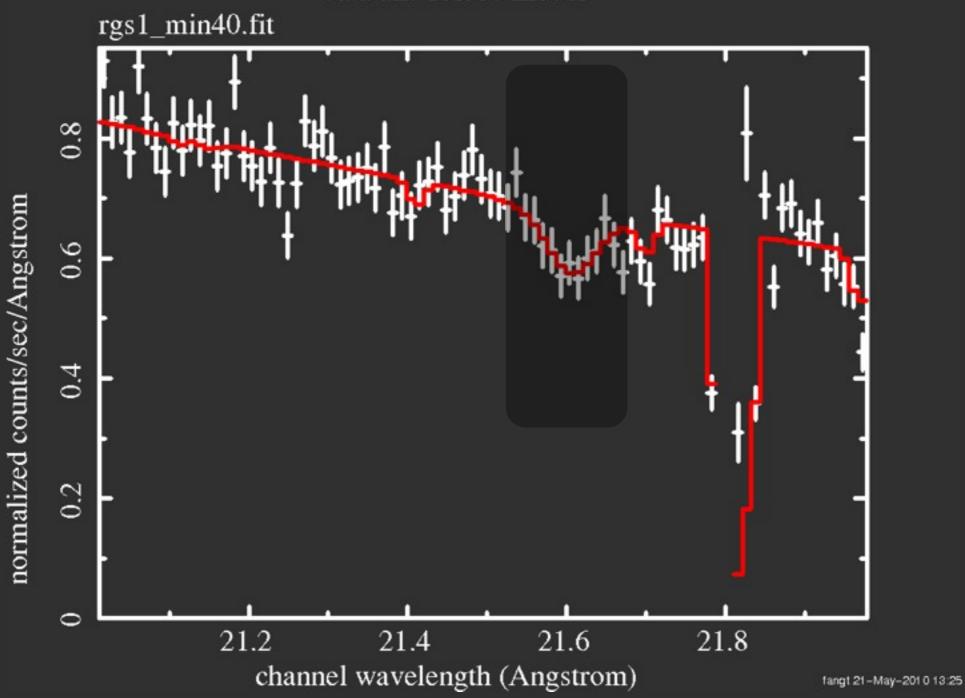
#### CYG X-2, CHANDRA

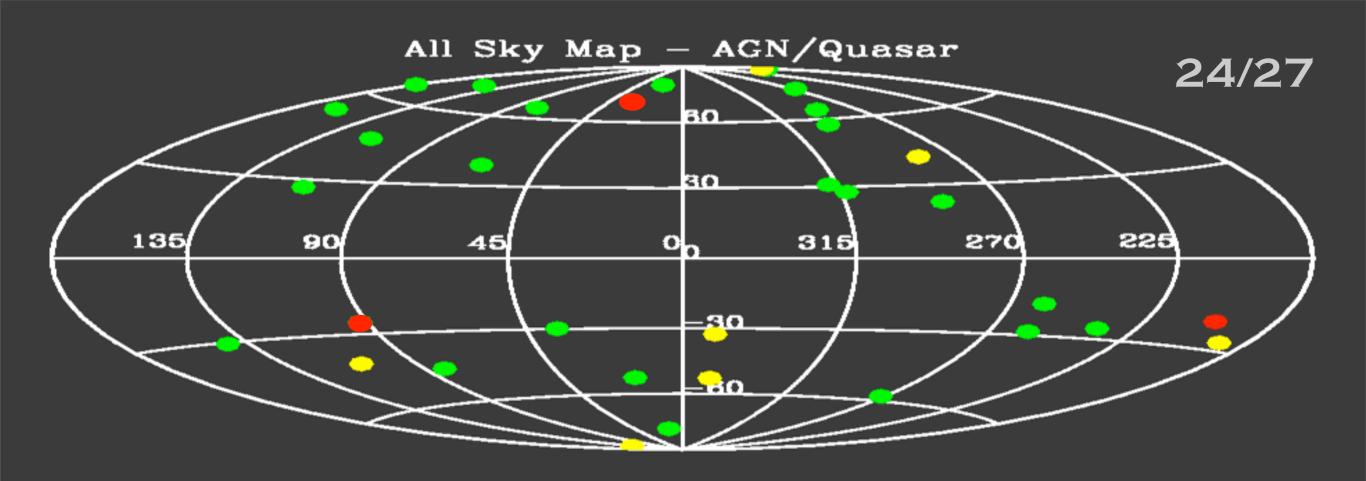


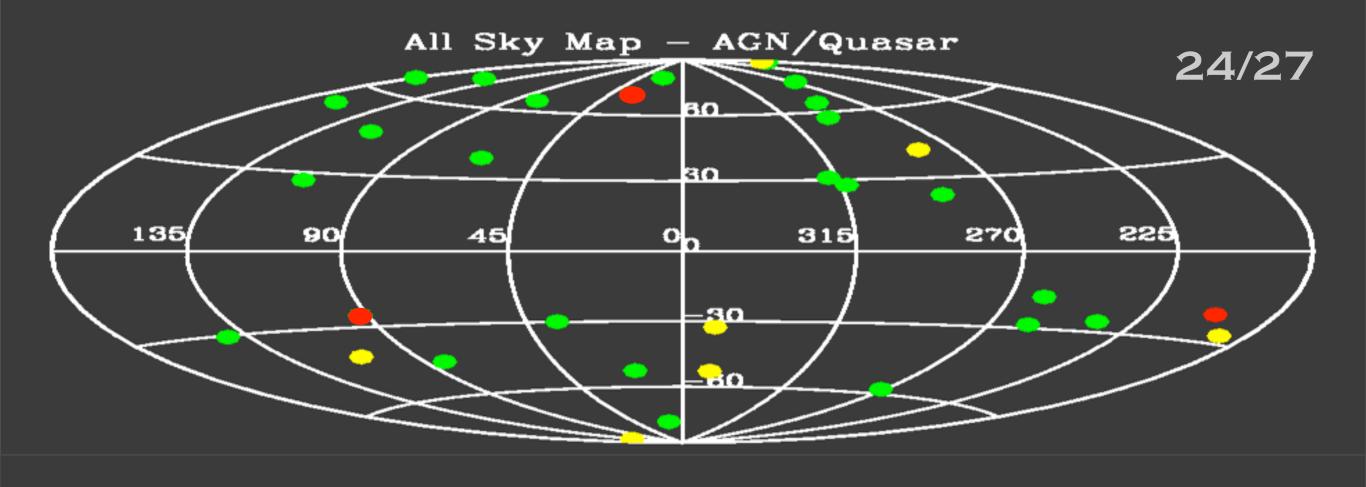
#### YAO ET AL. (2009)

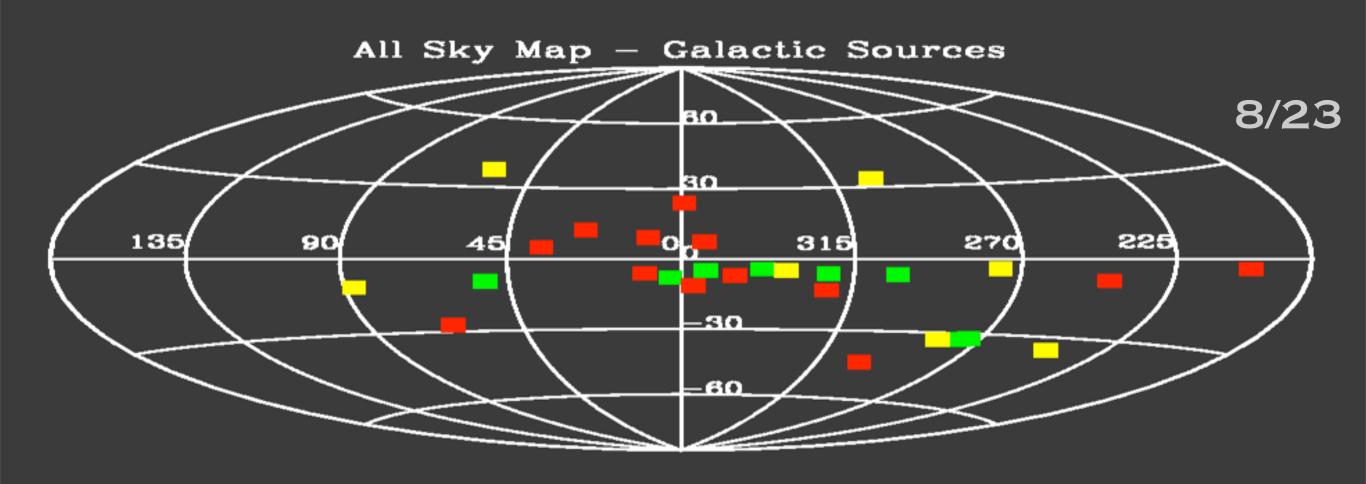
#### CYG X-2, XMM RGS

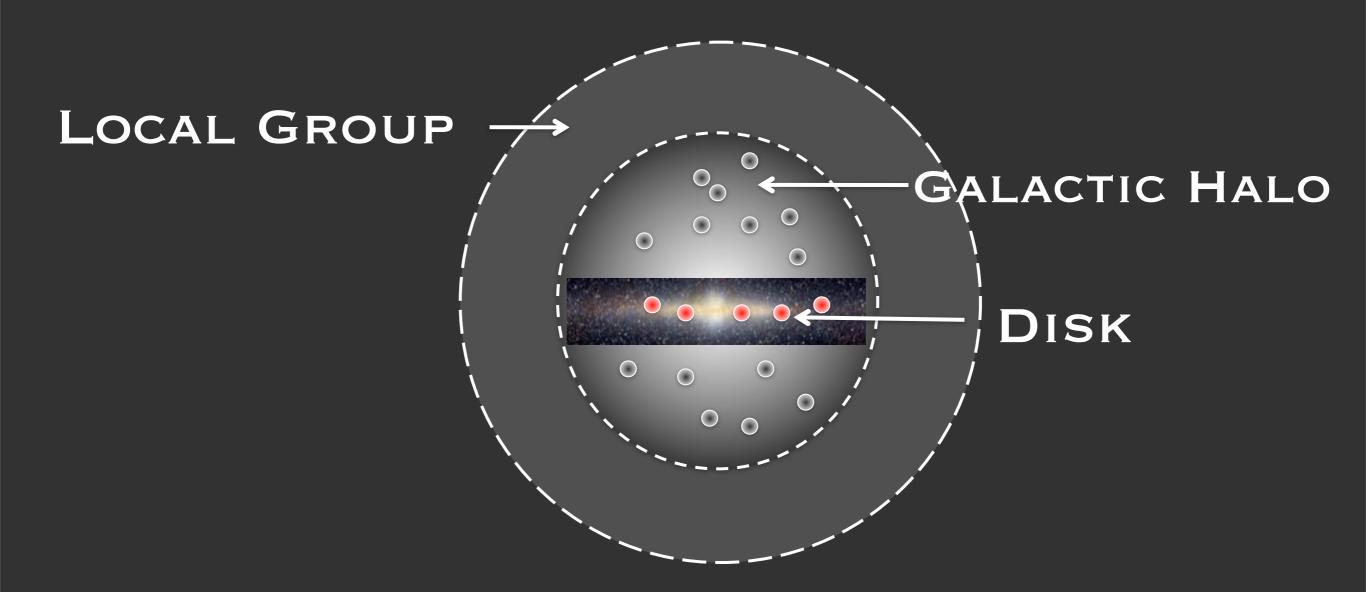


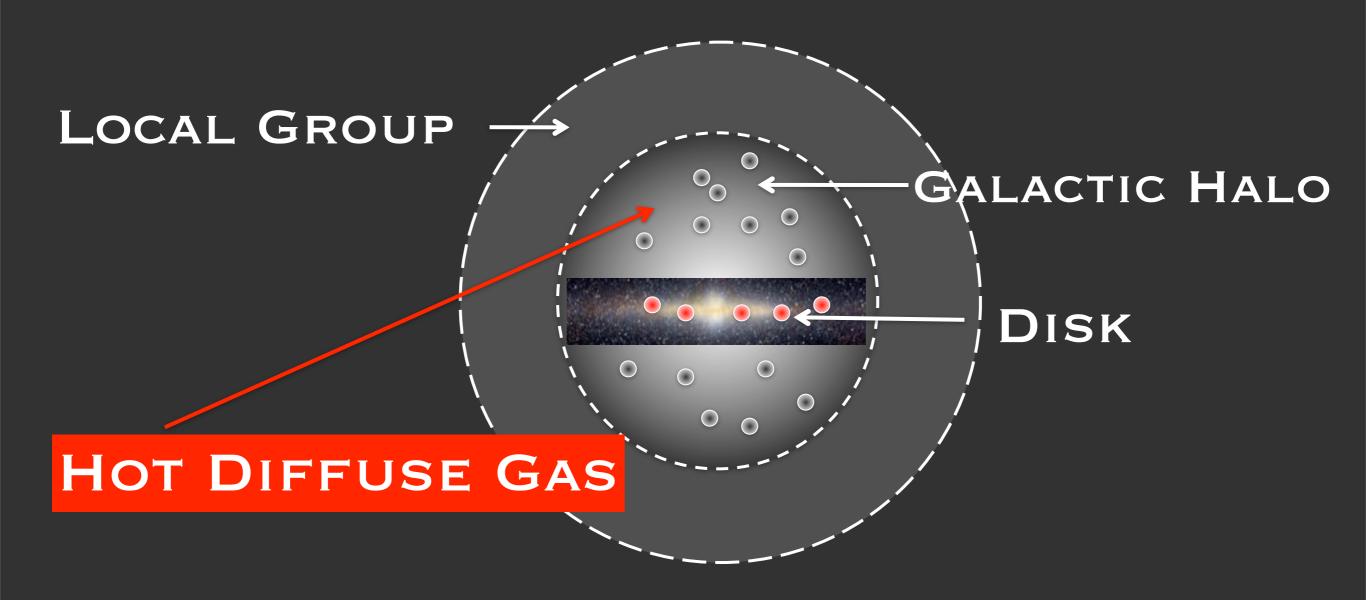


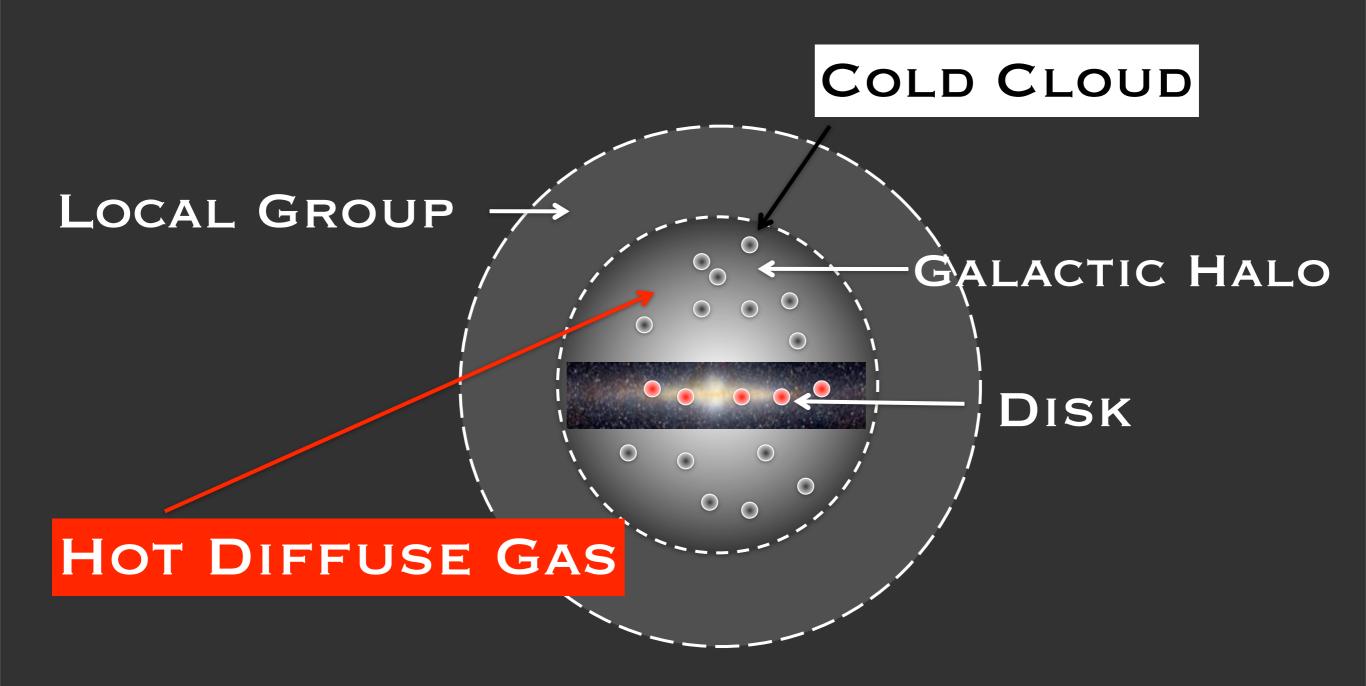


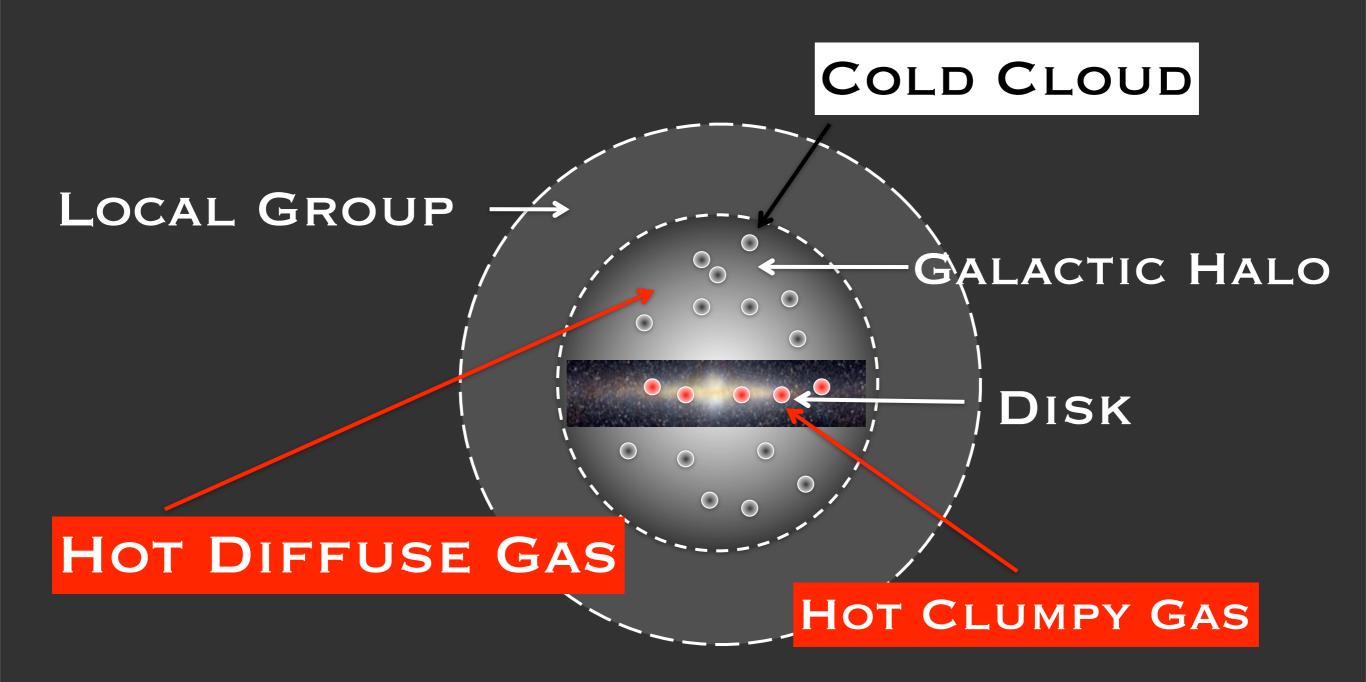












#### SUMMARY

- Searching for the missing baryons in the known structures
- Missing baryons: large-scale IGM, or galactic halos?
- International X-ray Observatory!